

# 1. Special Relativity (SR) — Content

## Content

- Introduction
  - Galilean Transformations
  - Axioms of Special Relativity
- Vectors, Tensors, and notation
- Invariants
- Lorentz transformations

## Links

- Lecture notes by David Hogg: <http://cosmo.nyu.edu/hogg/sr/sr.pdf>
  - or: <http://web.vu.lt/ff/t.gajdosik/files/2014/02/sr.pdf>

# 1. Special Relativity (SR) — Introduction

Galilean Invariance / Galilean transformations:  $t \rightarrow t'$ ,  $\vec{x} \rightarrow \vec{x}'$

Two inertial observers,  $O$  and  $O'$ ,

- measure the same absolute time (i.e.: 1 second = 1 second').
  - Time translations :  $t' = t + \tau$ ,  $\vec{x}' = \vec{x}$   
in index notation:  $t' = t + \tau$ ,  $x'_j = x_j$
- have at  $t = 0$  a relative distance  $\Delta\vec{r}$ .
  - Spatial translations :  $t' = t$ ,  $\vec{x}' = \vec{x} + \Delta\vec{r}$   
in index notation:  $t' = t$ ,  $x'_j = x_j + \Delta r_j$
- have coordinate systems that are rotated by a relative rotation  $\mathbf{R}$ .
  - Rotations :  $t' = t$ ,  $\vec{x}' = \mathbf{R} \cdot \vec{x}$ , where  $\mathbf{R}$  is an orthogonal matrix  
in index notation:  $t' = t$ ,  $x'_j = \mathbf{R}_{jk}x_k = \sum_{k=1}^3 \mathbf{R}_{jk}x_k$
- have a constant relative velocity  $\vec{v}$ , which can be zero, too.
  - Boosts :  $t' = t$ ,  $\vec{x}' = \vec{x} + \vec{v}t$   
in index notation:  $t' = t$ ,  $x'_j = x_j + v_j t$

# 1. Special Relativity (SR) — Introduction

## Galilean Group

- How Galilean transformations act on a quantum mechanical state.
- What is a **group**?
  - a **set** with a **binary operation**:
  - an example is the set of numbers  $\{0, 1, 2\}$  with the addition modulo 3 (i.e. taking only the remainder of the division by 3).
- Properties of a **group**
  - different transformations in the group do not give something that is outside the group.
  - two transformations in different order give either zero or another transformation.
- Each transformation depends on continuous parameters
  - The **Galilean Group** is a **Lie Group**.

# 1. Special Relativity (SR) — Introduction

## What's wrong with Galilean Invariance?

- Maxwell's equations describe the propagation of light depending on the electric permittivity and the magnetic permeability of the vacuum.
- If the vacuum is the same for every inertial observer, he has to measure the same speed of light regardless, who emitted it.
  - This is Einsteins second assumption!
- But then the addition of velocities described by the Galilean transformations are wrong.
- Lorentz transformations describe correctly the measurements done regarding the speed of light.
- Lorentz transformations include a transformation of the time, that the inertial observers measure.
- Absolut time is a concept, that is not able to describe nature.
  - That's wrong with the Galilean Invariance!

# 1. Special Relativity (SR) — Introduction

## Axioms of Special Relativity

- Every physical theory should look the same mathematically to every inertial observer.
- The speed of light in vacuum is independent from the movement of its emitting body.

## Consequences

- The speed of light in vacuum is maximum speed for any information.
- The world has to be described by a 4D space-time: Minkowski space.
- The simplest object is a scalar (field):  $\phi(x)$   
no structure except position and momentum.
- The next simplest object is a spinor (field):  $\psi^\alpha(x)$   
a vector (field) can be described as a double-spinor.

# 1. Special Relativity (SR) — Vectors, Tensors, and notation

the plane — i.e. 2D (Euclidean) space

- we can pick a **coordinate system** and describe points with coordinates
  - Cartesian coordinates  $(x, y)$
  - Polar coordinates  $(r, \theta)$
- a vector can be understood as a difference of points
  - position vector: difference between the position and the origin
- we can write the vector  $\vec{v}$ 
  - as a row  $(v_x, v_y)$  ... or as a column  $\begin{pmatrix} v_x \\ v_y \end{pmatrix}$
  - in index notation  $v_i$  or  $v^i$ , where we identify  $v_x = v_1$  and  $v_y = v_2$

We can understand the plane as being generated by **two vectors** :

$$\text{plane} = \text{point} + a\hat{x} + b\hat{y} \quad \text{with } a, b \in \mathcal{R}$$

- $\hat{x}$  and  $\hat{y}$  are said to **span** the plane, which is a **vectorspace**

# 1. Special Relativity (SR) — Vectors, Tensors, and notation

## multiplying vectors

- with a number, not a problem:  $c * \vec{a} = (c * a_x, c * a_y)$
- with another vector: what do we want to get?
  - a number  $\Rightarrow$  scalar product:  $\vec{a} \cdot \vec{b} := a_x * b_x + a_y * b_y$
  - another vector: there is no unique prescription ...
  - a tensor  $\Rightarrow$  tensor product:  $\vec{a} \otimes \vec{b}$ 
    - \* in index notation:  $a_j \otimes b_k = a_j b_k = (a \otimes b)_{jk}$

Geometric Algebra defines the geometric product of vectors :

for vectors  $a = \vec{a}$ ,  $b = \vec{b}$ , etc.

- $ab$  has a symmetric and an antisymmetric part:  $ab = a \cdot b + a \wedge b$ 
  - $aa = a \cdot a + a \wedge a = a^2$  is a number  
(the normal scalar product of the vector with itself)
  - $ba = b \cdot a + b \wedge a = a \cdot b - a \wedge b$  is a number plus a bivector
- $C = ab$  is called a multivector
- a multivector is NOT a tensor !

# 1. Special Relativity (SR) — Vectors, Tensors, and notation

## what is a tensor?

- an object that looks like the tensor product of vectors ...
  - it transforms like a tensor product of vectors would transform
- easiest imaginable in indexnotation:
  - a tensor is an object with indices  $t_{jkl}$  or  $t^{jkl}$  or  $t^j_{kl}$
- special tensors
  - a vector is a tensor of rank one: it has one index
  - a matrix is a tensor of rank two: it has two indices

## tensors form also a vectorspace

- multiplying a tensor with a number gives again a tensor
  - the resulting tensor is of the same dimensions as the initial one.
- adding tensors of the same dimension gives again a tensor



# 1. Special Relativity (SR) — Vectors, Tensors, and notation

## multiplying tensors

- one index of each can be treated like a scalar product
  - $\Rightarrow$  matrix multiplication
  - with  $a = a_{jk}$  and  $b = b_{mn}$ :  $(a \cdot b)_{jn} = \sum_k a_{jk} * b_{kn}$ 
    - \* here  $a$  and  $b$  can be understood as matrices
- in order to simplify the writing, we can omit the  $\sum$  symbol
  - $\Rightarrow$  **Einstein's summation convention**
  - one sums over repeated indices:  $a_{jk} * b_{kn} := \sum_k a_{jk} * b_{kn}$

index position can be used to distinguish objects

- example:
  - columnvector  $\vec{v} = v^i = \begin{pmatrix} v_x \\ v_y \end{pmatrix}$
  - rowvector  $(\vec{v})^\top = v_i = (v_x, v_y)$
- then a **matrix** has to have **upper and lower** index:  $a^j_k \neq a_k^j$

## 1. Special Relativity (SR) — Vectors, Tensors, and notation

in more dimensional space we just have more coordinates

- In 3D space (our 3D world):
  - $\vec{v} = (v_x, v_y, v_z) = v_i$  (in cartesian coordinates)
- In 4D Minkovsky space people do **not** write an arrow:
  - momentum  $p = (E = p^t, p^x, p^y, p^z) = (p^0, p^1, p^2, p^3) = p^\mu$ 
    - \* and the index is usually a greek letter:  $\mu, \nu, \rho$ , etc.
  - position  $r = (ct, x, y, z) = (x^0, x^1, x^2, x^3) = r^\mu$ 
    - \* time  $ct = x^0$  is measured like spacial distances in meter.
    - \* The constant speed of light  $c$  is used as the conversion factor between seconds and meters.

For the rest of the lecture we set  $c = 1$ . (i.e.:  $3 \cdot 10^8 \text{m} = 1\text{s}$ )

- so we measure time in seconds and distances in light-seconds
- or distances in meters and time in  $\sim 3$  "nanoseconds".

# 1. Special Relativity (SR) — Invariants

## What are invariant objects?

- Objects that are the same for every inertial observer.
- Examples in 3D: rotations or translations
  - the distances  $\ell$  between points:  $\ell^2 = (\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2$ .
  - the angle  $\alpha$  between directions:  $\cos \alpha = (\vec{a} \cdot \vec{b}) / (|\vec{a}| * |\vec{b}|)$ .
- In 4D Minkovsky space:  $(\Delta s)^2 = (\Delta t)^2 - (\Delta x)^2 - (\Delta y)^2 - (\Delta z)^2$ .
  - The time  $t$  is measured like spacial distances in meter.
  - The constant speed of light  $c$  is used as the conversion factor between seconds and meters.
- Any **scalar** product of four-vectors in Minkovsky space:

$$(p.q) = p^\mu q^\nu g_{\mu\nu} = p^0 q^0 - p^1 q^1 - p^2 q^2 - p^3 q^3 \ .$$

## 1. Special Relativity (SR) — Invariants

### What is the use of scalar products?

- Scalars are the same in every inertial frame.
  - If one knows its value in one frame, one knows it in every frame.
  - ⇒ Use the most comfortable frame to calculate the value of a scalar!
- Events  $A$  and  $B$  happen at a certain time in a certain place:
  - In every frame they can be described by four-vectors  $a^\mu = (a^0, a^1, a^2, a^3)$  and  $b^\mu = (b^0, b^1, b^2, b^3)$ .
  - Their relative position  $d^\mu = a^\mu - b^\mu$  is frame dependent.
  - But their "4-distance"  $d^2 = (d \cdot d)$  is invariant.
  - $d^2$  classifies the causal connection of  $A$  and  $B$ .

## 1. Special Relativity (SR) — Invariants

### Classification of $d^2$

- If  $d^2 > 0$  they are **time-like** separated:
  - one event happens before the other in every frame.
  - there is a frame, where  $A$  and  $B$  happen at the same position.
  - in this frame  $d^\mu = (\Delta t, 0, 0, 0)$  with  $\Delta t = \sqrt{d^2}$ .
- If  $d^2 = 0$  they are **light-like** related. If  $A \neq B$ :
  - there is no frame, where  $A$  and  $B$  happen at the same time.
  - there is no frame, where  $A$  and  $B$  happen at the same position.
  - there is a frame, where  $d^\mu = (\eta, \eta, 0, 0)$  with  $\eta$  arbitrary.
- If  $d^2 < 0$  they are **space-like** separated:
  - there is a frame, where  $A$  and  $B$  happen at the same time.
  - in this frame  $d^\mu = (0, \Delta s, 0, 0)$ , with  $\Delta s = \sqrt{-d^2}$ ,  
if the  $x$ -axis is oriented in the direction  $\overline{AB}$ .

## 1. Special Relativity (SR) — Invariants

### A special scalar product

- Particles are usually described by their energy-momentum four-vector:

$$p^\mu = (p^0, p^1, p^2, p^3) = (E, p_x, p_y, p_z) = (E, \vec{p})$$

- The mass of the particle is defined in its rest-frame:  $\vec{p} = 0$ .
- There, the energy-momentum four-vector is  $p^\mu = (m, 0)$ .
- Since  $p^2 = (p \cdot p)$  is a scalar, it is the same in every frame.
- In the rest-frame  $p^2 = m^2$ .
- Therefore in every frame

$$m^2 = E^2 - \vec{p}^2 \quad !$$

- This can be applied to collisions, too:  $(p_1 + p_2)^2$  is constant.

# 1. Special Relativity (SR) — Lorentz transformations

## Lorentz transformations

- relate the coordinate systems of two inertial observers.
- leave the "4-distance" invariant.
- assuming linearity, they can be written as

$$x'^{\mu} = \Lambda^{\mu}_{\nu} x^{\nu} + a^{\mu} .$$

- These are called inhomogeneous Lorentz transformations  $(\Lambda, a)$ .

Homogeneous Lorentz transformations have  $a^{\mu} = 0$ .

- They leave scalar products invariant:  $(p'.q') = (p.q)$ .
- They describe 3 Rotations and 3 Boosts (cf. the Galilean transformations).

## 1. Special Relativity (SR) — Lorentz transformations

Rotations are the same as in the Galilean transformations.

For Boosts between  $O$  and  $O'$  let us align the coordinate systems:

- The origins of  $O$  and  $O'$  should be at the same place at  $t = t' = 0$ .
- The constant relative velocity  $v$  between  $O$  and  $O'$  should point in the  $\hat{x}$ -direction for both,  $O$  and  $O'$ .
- $\hat{y}$  and  $\hat{z}$  should point in the same direction for both:  $y' = y$  and  $z' = z$ .
- Only  $ct = x^0$  and  $x = x^1$  are affected by such a boost:  
 $\Lambda^\mu{}_\nu = \delta^\mu{}_\nu$  for either  $\mu$  or  $\nu$  being 2 or 3.
- So with  $p' = \Lambda p$  and  $q' = \Lambda q$  we have  $(p'.q') - (p.q) = 0$ .
- Since  $y' = y$  and  $z' = z$  we can ignore  $\hat{y}$  and  $\hat{z}$  in the equation

$$0 = (p'.q') - (p.q) = (p'^0 q'^0 - p'^1 q'^1) - (p^0 q^0 - p^1 q^1) .$$



## 1. Special Relativity (SR) — Lorentz transformations

### Determining Boosts

$$\begin{aligned} 0 &= (\Lambda_0^0 p^0 + \Lambda_1^0 p^1)(\Lambda_0^0 q^0 + \Lambda_1^0 q^1) - (\Lambda_0^1 p^0 + \Lambda_1^1 p^1)(\Lambda_0^1 q^0 + \Lambda_1^1 q^1) \\ &\quad - (p^0 q^0 - p^1 q^1) \\ &= (\Lambda_0^0 \Lambda_0^0 - \Lambda_0^1 \Lambda_1^0 - 1)p^0 q^0 + (\Lambda_0^0 \Lambda_1^0 - \Lambda_0^1 \Lambda_1^1)p^0 q^1 \\ &\quad + (\Lambda_1^0 \Lambda_0^0 - \Lambda_1^1 \Lambda_1^0)p^1 q^0 + (\Lambda_1^0 \Lambda_1^0 - \Lambda_1^1 \Lambda_1^1 + 1)p^1 q^1 \end{aligned}$$

is solved by

$$\Lambda_0^0 = \Lambda_1^1 = \pm \cosh \eta \quad \Lambda_1^0 = \Lambda_0^1 = \mp \sinh \eta ,$$

where  $\eta$  is the "rapidity" of the boost. The usual choice is the upper sign.

How can we relate  $\eta$  to the relative velocity  $v$  between  $O$  and  $O'$ ?

- Let us take two events and describe them in  $O$  and  $O'$ :
  - $A$ : the origins of  $O$  and  $O'$  overlap; set  $t = t' = 0$ .
  - $B$ : at the origin of  $O'$  after the time  $t'$ , where  $t = \Delta t$ .

## 1. Special Relativity (SR) — Lorentz transformations

- The coordinates of  $A$  are  $a^\mu = a'^\mu = (0, 0, 0, 0)$ .
- The coordinates of  $B$ 
  - in  $O$  are  $b^\mu = (\Delta t, v\Delta t, 0, 0)$  because  $O'$  was moving with the constant relative velocity  $v$  for the time  $\Delta t$ .
  - in  $O'$  are  $b'^\mu = (t', 0, 0, 0)$  because  $B$  is at the origin of  $O'$ .
- But  $b'^\mu = \Lambda^\mu{}_\nu b^\nu = (\cosh \eta \Delta t - \sinh \eta v \Delta t, -\sinh \eta \Delta t + \cosh \eta v \Delta t, 0, 0)$ .

Therefore

$$\begin{aligned} t' &= \cosh \eta \Delta t - \sinh \eta v \Delta t \\ 0 &= -\sinh \eta \Delta t + \cosh \eta v \Delta t \end{aligned}$$

or

$$v = \frac{\sinh \eta}{\cosh \eta} = \tanh \eta \sim \eta \quad \text{for } \eta \text{ small.}$$

## 1. Special Relativity (SR) — Lorentz transformations

### Lorentz transformations on vectors

- Each vector  $V^\mu$  can be understood as the distance of two events.
- Its transformation is the same as for events in different inertial frames:

$$V'^\mu = \Lambda^\mu{}_\nu V^\nu .$$

- Since  $(V.W)$  is a scalar,  $(V'.W') = (V.W)$  :

$$V'^\mu W'_\mu = \Lambda^\mu{}_\nu V^\nu W'_\mu = V^\nu W_\nu .$$

- So  $\Lambda^\mu{}_\nu W'_\mu = W_\nu$  or  $W'_\mu = (\Lambda^\mu{}_\nu)^{-1} W_\nu$
- What is now the inverse  $(\Lambda(v))^{-1}$  ?
  - Obviously it should be  $\Lambda(-v)$  .

## 1. Special Relativity (SR) — Lorentz transformations

### More on vectors, the metric, and Lorentz transformations

- We defined the scalar product of **contravariant** vectors:

$$(p.q) = p^\mu q^\nu g_{\mu\nu} = p^0 q^0 - p^1 q^1 - p^2 q^2 - p^3 q^3 ,$$

where  $g_{\mu\nu} = g_{\nu\mu}$  is the metric with  $g_{00} = 1$ ,  $g_{ii} = -1$ , and  $g_{\mu \neq \nu} = 0$ .

- We can define **covariant** vectors with the index down:  $V_\mu = g_{\mu\nu} V^\nu$ .
- The index can be raised again by  $V^\mu = g^{\mu\nu} V_\nu$ .
- This obviously gives  $g^{\mu\nu} g_{\nu\rho} = g^{\nu\mu} g_{\nu\rho} = g^{\mu\nu} g_{\rho\nu} = \delta_\rho^\mu$ .
- That means for the Lorentz transformations:

$$V'_\mu = g_{\mu\lambda} V'^\lambda = g_{\mu\lambda} \Lambda^\lambda{}_\kappa V^\kappa = g_{\mu\lambda} \Lambda^\lambda{}_\kappa g^{\kappa\nu} V_\nu = (\Lambda^\mu{}_\nu)^{-1} V_\nu$$

or

$$(\Lambda^\mu{}_\nu)^{-1} = g_{\mu\lambda} \Lambda^\lambda{}_\kappa g^{\kappa\nu} = \Lambda_\mu{}^\nu .$$

# 1. Special Relativity (SR) — Lorentz transformations

## Lorentz transformations of fields

- Two observers,  $O$  and  $O'$ , can agree on a space-time point  $x$  by calling it an event  $X$ .
  - $X$  might have different coordinates  $x^\mu$  and  $x'^\mu$  in  $O$  and  $O'$ , but it is nevertheless the same point.
  - $O$  and  $O'$  can compare the value of different fields at that point  $X$ .

- The simplest field is the scalar field  $\phi(x)$ :

$$\phi'(X) = \phi(X) \ .$$

- The vector fields  $a^\mu(x)$  or  $a_\mu(x)$  transform like vectors:

$$a'^\mu(X) = \Lambda^\mu{}_\nu a^\nu(X) \quad a'_\mu(X) = \Lambda_\mu{}^\nu a_\nu(X) \ .$$

- Tensor fields  $t^{\mu\nu}_{\rho\kappa\lambda}(x)$  transform like the product of vectors:

$$t'^{\mu\nu}_{\rho\kappa\lambda}(X) = \Lambda^\mu{}_\alpha \Lambda^\nu{}_\beta \Lambda_\rho{}^\gamma \Lambda_\kappa{}^\delta \Lambda_\lambda{}^\epsilon t^{\alpha\beta}_{\gamma\delta\epsilon}(X) \ .$$