Reminder: 1. Special Relativity (SR) — four-vectors — slide 5

Vectors, Tensors, and **notation**

in the plane — i.e. in the 2D (Euclidean) space

- we can pick a coordinate system and describe points with coordinates
 - Cartesian coordinates (x, y)
 - Polar coordinates (r, θ)
- a vector can be understood as a difference of points
- position vector: difference between the position and the origin
- we can write the vector $ec{v}$
 - as a row (v_x, v_y)
 - or as a column $\begin{pmatrix} v_x \\ v_y \end{pmatrix}$

- or in index notation v_i or v^i , where we identify $v_x = v_1$ and $v_y = v_2$

Reminder: 1. Special Relativity (SR) — four-vectors — slide 8 Vectors, Tensors, and **notation**

in more dimensional space we just have more coordinates

• In 3D space (our 3D world):

 $-\vec{v} = (v_x, v_y, v_z) = v_i$ (in cartesian coordinates)

- In **4D** Minkovsky space people do not write an arrow:
 - momentum $p = (E = p^t, p^x, p^y, p^z) = (p^0, p^1, p^2, p^3) = p^{\mu}$
 - * and the index is usually a greek letter: μ , ν , ρ , etc.
 - position $r = (ct, x, y, z) = (x^0, x^1, x^2, x^3) = r^{\mu}$
 - * time $ct = x^0$ is measured like spacial distances in meters
 - \ast The constant speed of light c is used as the conversion factor between seconds and meters

For the rest of the lecture we set c = 1. (i.e.: $3 \cdot 10^8 \text{m} = 1$ s)

- so we measure time in seconds and distances in light-seconds (=300.000km)
- or distances in meters and time in "3 nanoseconds" (the time light needs to travel 1m)

Reminder: 1. Special Relativity (SR) — Invariants — slide 9

What are invariant objects?

- Objects that are the same for every inertial observer
- Examples in 3D: rotations or translations
 - the distances ℓ between points: $\ell^2 = (\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2$
 - the angle α between directions: $\cos \alpha = (\vec{a} \cdot \vec{b})/(|\vec{a}| * |\vec{b}|)$
- In 4D Minkovsky space: $(\Delta s)^2 = (\Delta t)^2 (\Delta x)^2 (\Delta y)^2 (\Delta z)^2$
 - The time t is measured like spacial distances in meters
 - The constant speed of light \boldsymbol{c} is used as the conversion factor between seconds and meters

• Any scalar product of four-vectors in Minkovsky space:

$$(p.q) = p^{\mu}q^{\nu}g_{\mu\nu} = p^{0}q^{0} - p^{1}q^{1} - p^{2}q^{2} - p^{3}q^{3} \quad (1)$$

Reminder: 1. Special Relativity (SR) — characterisation — slide 11 Classification of $d^2 = (d \cdot d) = d^{\mu}d_{\mu} = d^{\mu}d^{\nu}g_{\mu\nu}$

- If $d^2 > 0$ they are **time-like** separated:
 - one event happens before the other in every frame
 - there is a frame, where A and B happen at the same position
 - in this frame $d^{\mu} = (\Delta t, 0, 0, 0)$ with $\Delta t = \sqrt{d^2}$
- If $d^2 = 0$ they are **light-like** related. If $A \neq B$:
 - there is no frame, where A and B happen at the same time
 - there is no frame, where A and B happen at the same position
 - there is a frame, where $d^{\mu} = (\eta, \eta, 0, 0)$ with η arbitrary
- If $d^2 < 0$ they are **space-like** separated:
 - there is a frame, where A and B happen at the same time
 - in this frame $d^{\mu} = (0, \Delta s, 0, 0)$, with $\Delta s = \sqrt{-d^2}$, if the *x*-axis is oriented in the direction \overline{AB}

Reminder: 1. Special Relativity (SR) — characterisation — slide 12 Special scalar products

• Particles are described by their energy-momentum four-vector:

$$p^{\mu} = (p^0, p^1, p^2, p^3) = (E, p_x, p_y, p_z) = (E, \vec{p})$$
 (2)

- The mass of the particle is defined in its rest-frame: $\vec{p} = 0$
- There, the energy-momentum four-vector is $p^{\mu} = (m, 0)$
- Since $p^2 = (p \cdot p)$ is a scalar, it is the same in every frame
- In the rest-frame $p^2 = m^2 \vec{0}^2 = m^2$
- Therefore in every frame

$$m^2 = E^2 - \vec{p}^2 \quad ! \tag{3}$$

- This can be applied to collisions, too: $(p_1 + p_2)^2$ is constant
 - In the rest-frame of $(p_1 + p_2)$ we have $\vec{p_1} + \vec{p_2} = 0 \Rightarrow (p_1 + p_2)^2 = (E_1 + E_2)^2$
 - * E_1 and E_2 are the energy values of p_1 and p_2 in the rest-frame of $(p_1 + p_2)!$

2. Special Relativity (SR) — Summary of SR1

Space-time objects are described by four-vectors

for better space management on the slides I write all vectors as row vectors

• position $r^{\mu} = (ct, x, y, z) = (x^0, x^1, x^2, x^3)$

– Time by the zero component $r^{\rm 0}=x^{\rm 0}=ct$

- Space by the components > 0
$$\vec{r} = r^i = x^i = (x^1, x^2, x^3) = (x, y, z)$$

- momentum $p^{\mu} = (E/c, p_x, p_y, p_z) = (p^0, p^1, p^2, p^3)$
 - Energy (or mass) by the zero component $p^0 = E/c$
 - three momentum by the components > 0 $\vec{p} = (p^1, p^2, p^3) = (p_x, p_y, p_z)$
- mass is defined in the rest-frame: $\vec{p} = 0$

- There, the energy-momentum four-vector is $p^{\mu} = (m, 0)$

- 2. Special Relativity (SR) raising and lowering indices The metric $g^{\mu\nu}$ (or $g_{\mu\nu}$) is used to raise (or lower) indices
 - in flat Minkovsky space the components of the metric tensor are

$$g_{00} = g^{00} = 1$$
 $g_{ii} = g^{ii} = -1$ $g_{j \neq k} = g^{j \neq k} = 0$ (4)

• that gives for the contravariant four vector $p^{\mu} = (E/c, p_x, p_y, p_z)$

- the covariant four vector

$$p_{\mu} = g_{\mu\nu}p^{\nu} = g_{\mu0}E/c + g_{\mu1}p_x + g_{\mu2}p_y + g_{\mu3}p_z \tag{5}$$

SO

$$p_{0} = g_{00}E/c + g_{01}p_{x} + g_{02}p_{y} + g_{03}p_{z} = 1 \times E/c + 0 \times p_{x} + 0 \times p_{y} + 0 \times p_{z} = E/c$$

$$p_{1} = g_{10}E/c + g_{11}p_{x} + g_{12}p_{y} + g_{13}p_{z} = 0 \times E/c + (-1) \times p_{x} + 0 \times p_{y} + 0 \times p_{z} = -p_{x}$$

$$p_{2} = g_{20}E/c + g_{21}p_{x} + g_{22}p_{y} + g_{23}p_{z} = \dots = -p_{y}$$

$$p_{3} = g_{30}E/c + g_{31}p_{x} + g_{32}p_{y} + g_{33}p_{z} = \dots = -p_{z}$$

and hence

$$p_{\mu} = (E/c, -p_x, -p_y, -p_z)$$
 (6)

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2. Special Relativity (SR) — raising and lowering indices

The metric $g^{\mu\nu}$ (or $g_{\mu\nu}$) is used to raise (or lower) indices

• check yourself that the invariant product, eq.(1), is really invariant

$$(p.q) = p^{\mu}q^{\nu}g_{\mu\nu} = p^{\mu}q_{\mu} = p_{\nu}q^{\nu} = p_{\mu}q_{\nu}g^{\mu\nu}$$
(7)

- for tensors we have to operate on each index
 - lowering both indices of a tensor $s^{\mu\nu}$

$$s_{\mu\nu} = g_{\mu\alpha}g_{\nu\beta}s^{\alpha\beta} = g_{\mu\mu'}g_{\nu\nu'}s^{\mu'\nu'} \tag{8}$$

- what we **cannot** write:

$$s_{\mu\nu} = g_{\mu\mu}g_{\mu\nu}s^{\mu\nu} \tag{9}$$

- we would make several mistakes:

* the indices left and right do not match: $_{\mu\nu}\neq ^{\mu\nu}$

* the summation is not defined: we have 4 appearances of μ or ν

2. Special Relativity (SR) — Lorentz γ -factor

Lorentz transformations

... will be derived in two weeks

$$ct' = \gamma(ct - \beta x)$$
 $x' = \gamma(x - \beta ct)$ (10)

where

$$\beta := v/c$$
 (11)
 $\gamma := [1 - \beta^2]^{-1/2}$ (12)

- Time dilation is described in chapter 2.1 of D.Hogg's notes
 - by means of the thought experiment of a light clock
 - it can be summarized by

moving clocks go slower by the factor γ

- that means: the decay time of a moving particle is longer: $t=\gamma\tau$
- Length contraction is discussed in a similar way in chapter 2.3:

a moving object seems shorter by the factor γ

- the length of a moving volume is shorter: $\ell_{\text{measured}} = \frac{1}{\gamma} \ell_{\text{rest-frame}}$