

1. Special Relativity (SR) — Content

Content

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Links

- Lecture notes by David Hogg: <http://cosmo.nyu.edu/hogg/sr/sr.pdf>
 - or: <http://www.tfk.ff.vu.lt/~garfield/WoP/sr.pdf>

1. Special Relativity (SR) — Introduction

Galilean Invariance / Galilean transformations: $t \rightarrow t'$, $\vec{x} \rightarrow \vec{x}'$

Two inertial observers, O and O' ,

- measure the same absolute time (i.e.: 1 second = 1 second').
 - Time translations : $t' = t + \tau$, $\vec{x}' = \vec{x}$
in index notation: $t' = t + \tau$, $x'_j = x_j$
- have at $t = 0$ a relative distance $\Delta\vec{r}$.
 - Spatial translations : $t' = t$, $\vec{x}' = \vec{x} + \Delta\vec{r}$
in index notation: $t' = t$, $x'_j = x_j + \Delta r_j$
- have coordinate systems that are rotated by a relative rotation \mathbf{R} .
 - Rotations : $t' = t$, $\vec{x}' = \mathbf{R} \cdot \vec{x}$, where \mathbf{R} is an orthogonal matrix
in index notation: $t' = t$, $x'_j = \mathbf{R}_{jk}x_k = \sum_{k=1}^3 \mathbf{R}_{jk}x_k$
- have a constant relative velocity \vec{v} , which can be zero, too.
 - Boosts : $t' = t$, $\vec{x}' = \vec{x} + \vec{v}t$
in index notation: $t' = t$, $x'_j = x_j + v_j t$

1. Special Relativity (SR) — Introduction

Galilean Group

- How the Galilean transformations act on a quantum mechanical state.
- What is a **group**?
 - a **set** with a **binary operation**:
 - an example is the set of numbers $\{0, 1, 2\}$ with the addition modulo 3 (i.e. taking only the remainder of the division by 3).
- Properties of a **group**
 - different transformations in the group do not give something that is outside the group.
 - two transformations in different order give either zero or another transformation.
- Each transformation depends on continuous parameters
 - The **Galilean Group** is a **Lie Group**.

1. Special Relativity (SR) — Introduction

What's wrong with Galilean Invariance?

- Maxwell's equations describe the propagation of light depending on the electric permittivity and the magnetic permeability of the vacuum.
- If the vacuum is the same for every inertial observer, he has to measure the same speed of light regardless, who emitted it.
 - This is Einsteins second assumption!
- But then the addition of velocities described by the Galilean transformations are wrong.
- Lorentz transformations describe correctly the measurements done regarding the speed of light.
- Lorentz transformations include a transformation of the time, that the inertial observers measure.
- Absolut time is a concept, that is not able to describe nature.
 - That's wrong with the Galilean Invariance!

1. Special Relativity (SR) — Introduction

Axioms of Special Relativity

- Every physical theory should look the same mathematically to every inertial observer.
- The speed of light in vacuum is independent from the movement of its emitting body.

Consequences

- The speed of light in vacuum is maximum speed for any information.
- The world has to be described by a 4D space-time: Minkowski space.
- The simplest object is a scalar (field): $\phi(x)$
no structure except position and momentum.
- The next simplest object is a spinor (field): $\psi^\alpha(x)$
a vector (field) can be described as a double-spinor.

1. Special Relativity (SR) — Vectors, Tensors, and notation

the plane — i.e. 2D (Euclidean) space

- we can pick a **coordinate system** and describe points with coordinates
 - Cartesian coordinates (x, y)
 - Polar coordinates (r, θ)
- a vector can be understood as a difference of points
 - position vector: difference between the position and the origin
- we can write the vector \vec{v}
 - as a row (v_x, v_y) ... or as a column $\begin{pmatrix} v_x \\ v_y \end{pmatrix}$
 - in index notation v_i or v^i , where we identify $v_x = v_1$ and $v_y = v_2$

We can understand the plane as being generated by **two vectors** :

$$\text{plane} = \text{point} + a\hat{x} + b\hat{y} \quad \text{with } a, b \in \mathcal{R}$$

- \hat{x} and \hat{y} are said to **span** the plane, which is a **vectorspace**

1. Special Relativity (SR) — Vectors, Tensors, and notation

multiplying vectors

- with a number, not a problem: $c * \vec{a} = (c * a_x, c * a_y)$
- with another vector: what do we want to get?
 - a number \Rightarrow scalar product: $\vec{a} \cdot \vec{b} := a_x * b_x + a_y * b_y$
 - another vector: there is no unique prescription ...
 - a tensor \Rightarrow tensor product: $\vec{a} \otimes \vec{b}$
 - * in index notation: $a_j \otimes b_k = a_j b_k = (a \otimes b)_{jk}$

Geometric Algebra defines the geometric product of vectors :

for vectors $a = \vec{a}$, $b = \vec{b}$, etc.

- ab has a symmetric and an antisymmetric part: $ab = a \cdot b + a \wedge b$
 - $aa = a \cdot a + a \wedge a = a^2$ is a number
(the normal scalar product of the vector with itself)
 - $ba = b \cdot a + b \wedge a = a \cdot b - a \wedge b$ is a number plus a bivector
- $C = ab$ is called a multivector
- a multivector is NOT a tensor !

1. Special Relativity (SR) — Vectors, Tensors, and notation

what is a tensor?

- an object that looks like the tensor product of vectors ...
 - it transforms like a tensor product of vectors would transform
- easiest imaginable in indexnotation:
 - a tensor is an object with indices t_{jkl} or t^{jkl} or t^j_{kl}
- special tensors
 - a vector is a tensor of rank one: it has one index
 - a matrix is a tensor of rank two: it has two indices

tensors form also a vectorspace

- multiplying a tensor with a number gives again a tensor
 - the resulting tensor is of the same dimensions as the initial one.
- adding tensors of the same dimension gives again a tensor

1. Special Relativity (SR) — Vectors, Tensors, and notation

multiplying tensors

- one index of each can be treated like a scalar product
 - ⇒ matrix multiplication
 - with $a = a_{jk}$ and $b = b_{mn}$: $(a \cdot b)_{jn} = \sum_k a_{jk} * b_{kn}$
 - * here a and b can be understood as matrices
- in order to simplify the writing, we can omit the \sum symbol
 - ⇒ **Einsteins summation convention**
 - one sums over repeated indices: $a_{jk} * b_{kn} := \sum_k a_{jk} * b_{kn}$

index position can be used to distinguish objects

- example:
 - columnvector $\vec{v} = v^i = \begin{pmatrix} v_x \\ v_y \end{pmatrix}$
 - rowvector $(\vec{v})^\top = v_i = (v_x, v_y)$
- then a **matrix** has to have **upper and lower** index: $a^j_k \neq a_k^j$

1. Special Relativity (SR) — Vectors, Tensors, and notation

in more dimensional space we just have more coordinates

- In 3D space (our 3D world):
 - $\vec{v} = (v_x, v_y, v_z) = v_i$ (in cartesian coordinates)
- In 4D Minkovsky space people do **not** write an arrow:
 - momentum $p = (E = p^t, p^x, p^y, p^z) = (p^0, p^1, p^2, p^3) = p^\mu$
 - * and the index is usually a greek letter: μ, ν, ρ , etc.
 - position $r = (ct, x, y, z) = (x^0, x^1, x^2, x^3) = r^\mu$
 - * time $ct = x^0$ is measured like spacial distances in meter.
 - * The constant speed of light c is used as the conversion factor between seconds and meters.

For the rest of the lecture we set $c = 1$. (i.e.: $3 \cdot 10^8 \text{m} = 1\text{s}$)

- so we measure time in seconds and distances in light-seconds
- or distances in meters and time in 3 nanometers.

1. Special Relativity (SR) — Invariants

What are invariant objects?

- Objects that are the same for every inertial observer.
- Examples in 3D: rotations or translations
 - the distances ℓ between points: $\ell^2 = (\Delta x)^2 + (\Delta y)^2 + (\Delta z)^2$.
 - the angle α between directions: $\cos \alpha = (\vec{a} \cdot \vec{b}) / (|\vec{a}| * |\vec{b}|)$.
- In 4D Minkovsky space: $(\Delta s)^2 = (\Delta t)^2 - (\Delta x)^2 - (\Delta y)^2 - (\Delta z)^2$.
 - The time t is measured like spacial distances in meter.
 - The constant speed of light c is used as the conversion factor between seconds and meters.
- Any **scalar** product of four-vectors in Minkovsky space:

$$(p.q) = p^\mu q^\nu g_{\mu\nu} = p^0 q^0 - p^1 q^1 - p^2 q^2 - p^3 q^3 .$$

1. Special Relativity (SR) — Invariants

What is the use of scalar products?

- Scalars are the same in every inertial frame.
 - If one knows its value in one frame, one knows it in every frame.
 - ⇒ Use the most comfortable frame to calculate the value of a scalar!
- Events A and B happen at a certain time in a certain place:
 - In every frame they can be described by four-vectors $a^\mu = (a^0, a^1, a^2, a^3)$ and $b^\mu = (b^0, b^1, b^2, b^3)$.
 - Their relative position $d^\mu = a^\mu - b^\mu$ is frame dependent.
 - But their "4-distance" $d^2 = (d \cdot d)$ is invariant.
 - d^2 classifies the causal connection of A and B .

1. Special Relativity (SR) — Invariants

Classification of d^2

- If $d^2 > 0$ they are **time-like** separated:
 - one event happens before the other in every frame.
 - there is a frame, where A and B happen at the same position.
 - in this frame $d^\mu = (\Delta t, 0, 0, 0)$ with $\Delta t = \sqrt{d^2}$.
- If $d^2 = 0$ they are **light-like** related. If $A \neq B$:
 - there is no frame, where A and B happen at the same time.
 - there is no frame, where A and B happen at the same position.
 - there is a frame, where $d^\mu = (\eta, \eta, 0, 0)$ with η arbitrary.
- If $d^2 < 0$ they are **space-like** separated:
 - there is a frame, where A and B happen at the same time.
 - in this frame $d^\mu = (0, \Delta s, 0, 0)$, with $\Delta s = \sqrt{-d^2}$,
if the x -axis is oriented in the direction \overline{AB} .

1. Special Relativity (SR) — Invariants

A special scalar product

- Particles are usually described by their energy-momentum four-vector:

$$p^\mu = (p^0, p^1, p^2, p^3) = (E, p_x, p_y, p_z) = (E, \vec{p})$$

- The mass of the particle is defined in its rest-frame: $\vec{p} = 0$.
- There, the energy-momentum four-vector is $p^\mu = (m, 0)$.
- Since $p^2 = (p \cdot p)$ is a scalar, it is the same in every frame.
- In the rest-frame $p^2 = m^2$.
- Therefore in every frame

$$m^2 = E^2 - \vec{p}^2 \quad !$$

- This can be applied to collisions, too: $(p_1 + p_2)^2$ is constant.

1. Special Relativity (SR) — Lorentz transformations

Lorentz transformations

- relate the coordinate systems of two inertial observers.
- leave the "4-distance" invariant.
- assuming linearity, they can be written as

$$x'^{\mu} = \Lambda^{\mu}_{\nu} x^{\nu} + a^{\mu} .$$

- These are called **inhomogeneous Lorentz transformations** (Λ, a) .

Homogeneous Lorentz transformations have $a^{\mu} = 0$.

- They leave **scalar products** invariant: $(p'.q') = (p.q)$.
- They describe **3 Rotations** and **3 Boosts**
(cf. the Galilean transformations).

1. Special Relativity (SR) — Lorentz transformations

Rotations are the same as in the Galilean transformations.

For Boosts between O and O' let us align the coordinate systems:

- The origins of O and O' should be at the same place at $t = t' = 0$.
- The constant relative velocity v between O and O' should point in the \hat{x} -direction for both, O and O' .
- \hat{y} and \hat{z} should point in the same direction for both: $y' = y$ and $z' = z$.
- Only $ct = x^0$ and $x = x^1$ are affected by such a boost:
 $\Lambda^\mu_\nu = \delta^\mu_\nu$ for either μ or ν being 2 or 3.
- So with $p' = \Lambda p$ and $q' = \Lambda q$ we have $(p'.q') - (p.q) = 0$.
- Since $y' = y$ and $z' = z$ we can ignore \hat{y} and \hat{z} in the equation

$$0 = (p'.q') - (p.q) = (p'^0 q'^0 - p'^1 q'^1) - (p^0 q^0 - p^1 q^1) .$$

1. Special Relativity (SR) — Lorentz transformations

Determining Boosts

$$\begin{aligned} 0 &= (\Lambda_0^0 p^0 + \Lambda_1^0 p^1)(\Lambda_0^0 q^0 + \Lambda_1^0 q^1) - (\Lambda_0^1 p^0 + \Lambda_1^1 p^1)(\Lambda_0^1 q^0 + \Lambda_1^1 q^1) \\ &\quad - (p^0 q^0 - p^1 q^1) \\ &= (\Lambda_0^0 \Lambda_0^0 - \Lambda_0^1 \Lambda_1^0 - 1)p^0 q^0 + (\Lambda_0^0 \Lambda_1^0 - \Lambda_0^1 \Lambda_1^1)p^0 q^1 \\ &\quad + (\Lambda_1^0 \Lambda_0^0 - \Lambda_1^1 \Lambda_1^0)p^1 q^0 + (\Lambda_1^0 \Lambda_1^0 - \Lambda_1^1 \Lambda_1^1 + 1)p^1 q^1 \end{aligned}$$

is solved by

$$\Lambda_0^0 = \Lambda_1^1 = \pm \cosh \eta \quad \Lambda_1^0 = \Lambda_0^1 = \mp \sinh \eta ,$$

where η is the "rapidity" of the boost. The usual choice is the upper sign.

How can we relate η to the relative velocity v between O and O' ?

- Let us take two events and describe them in O and O' :
 - A : the origins of O and O' overlap; set $t = t' = 0$.
 - B : at the origin of O' after the time t' , where $t = \Delta t$.

1. Special Relativity (SR) — Lorentz transformations

- The coordinates of A are $a^\mu = a'^\mu = (0, 0, 0, 0)$.
- The coordinates of B
 - in O are $b^\mu = (\Delta t, v\Delta t, 0, 0)$ because O' was moving with the constant relative velocity v for the time Δt .
 - in O' are $b'^\mu = (t', 0, 0, 0)$ because B is at the origin of O' .
- But $b'^\mu = \Lambda^\mu{}_\nu b^\nu = (\cosh \eta \Delta t - \sinh \eta v \Delta t, -\sinh \eta \Delta t + \cosh \eta v \Delta t, 0, 0)$.
Therefore

$$\begin{aligned}t' &= \cosh \eta \Delta t - \sinh \eta v \Delta t \\0 &= -\sinh \eta \Delta t + \cosh \eta v \Delta t\end{aligned}$$

or

$$v = \frac{\sinh \eta}{\cosh \eta} = \tanh \eta \sim \eta \quad \text{for } \eta \text{ small.}$$

1. Special Relativity (SR) — Lorentz transformations

Lorentz transformations on vectors

- Each vector V^μ can be understood as the distance of two events.
- Its transformation is the same as for events in different inertial frames:

$$V'^\mu = \Lambda^\mu{}_\nu V^\nu .$$

- Since $(V.W)$ is a scalar, $(V'.W') = (V.W)$:

$$V'^\mu W'_\mu = \Lambda^\mu{}_\nu V^\nu W'_\mu = V^\nu W_\nu .$$

- So $\Lambda^\mu{}_\nu W'_\mu = W_\nu$ or $W'_\mu = (\Lambda^\mu{}_\nu)^{-1} W_\nu$
- What is now the inverse $(\Lambda(v))^{-1}$?
 - Obviously it should be $\Lambda(-v)$.

1. Special Relativity (SR) — Lorentz transformations

More on vectors, the metric, and Lorentz transformations

- We defined the scalar product of **contravariant** vectors:

$$(p.q) = p^\mu q^\nu g_{\mu\nu} = p^0 q^0 - p^1 q^1 - p^2 q^2 - p^3 q^3 ,$$

where $g_{\mu\nu} = g_{\nu\mu}$ is the metric with $g_{00} = 1$, $g_{ii} = -1$, and $g_{\mu\neq\nu} = 0$.

- We can define **covariant** vectors with the index down: $V_\mu = g_{\mu\nu} V^\nu$.
- The index can be raised again by $V^\mu = g^{\mu\nu} V_\nu$.
- This obviously gives $g^{\mu\nu} g_{\nu\rho} = g^{\nu\mu} g_{\nu\rho} = g^{\mu\nu} g_{\rho\nu} = \delta_\rho^\mu$.
- That means for the Lorentz transformations:

$$V'_\mu = g_{\mu\lambda} V'^\lambda = g_{\mu\lambda} \Lambda^\lambda{}_\kappa V^\kappa = g_{\mu\lambda} \Lambda^\lambda{}_\kappa g^{\kappa\nu} V_\nu = (\Lambda^\mu{}_\nu)^{-1} V_\nu$$

or

$$(\Lambda^\mu{}_\nu)^{-1} = g_{\mu\lambda} \Lambda^\lambda{}_\kappa g^{\kappa\nu} = \Lambda_\mu{}^\nu .$$

1. Special Relativity (SR) — Lorentz transformations

Lorentz transformations of fields

- Two observers, O and O' , can agree on a space-time point x by calling it an event X .
 - X might have different coordinates x^μ and x'^μ in O and O' , but it is nevertheless the same point.
 - O and O' can compare the value of different fields at that point X .

- The simplest field is the scalar field $\phi(x)$:

$$\phi'(X) = \phi(X) .$$

- The vector fields $a^\mu(x)$ or $a_\mu(x)$ transform like vectors:

$$a'^\mu(X) = \Lambda^\mu{}_\nu a^\nu(X) \quad a'_\mu(X) = \Lambda_\mu{}^\nu a_\nu(X) .$$

- Tensor fields $t^{\mu\nu}{}_{\rho\kappa\lambda}(x)$ transform like the product of vectors:

$$t'^{\mu\nu}{}_{\rho\kappa\lambda}(X) = \Lambda^\mu{}_\alpha \Lambda^\nu{}_\beta \Lambda_\rho{}^\gamma \Lambda_\kappa{}^\delta \Lambda_\lambda{}^\epsilon t^{\alpha\beta}{}_{\gamma\delta\epsilon}(X) .$$